

FISICA: FAR-INFRARED SPACE INTERFEROMETER CRITICAL ASSESSMENT.

SCIENCE DRIVERS DEFINITION
AND
TECHNOLOGY DEVELOPMENT

Giorgio Savini on behalf of the FP7-FISICA Consortium

THE CONSORTIUM



Nicola Baccichet, Roser Juanola-Parramon, **Giorgio Savini**, Bruce Swinyard, Amelie Guisseau



Peter Ade, Matt Griffin, Pete Hargrave, Georgina Klemencic, **Enzo Pascale**, Rashmi Sudiwala.



Rob Ivison → **Wayne Holland**, John Lightfoot



Martyn Jones , **David Walker**



Colm Bracken, Anthony Donohoe, **Anthony Murphy**, Creidhe O'Sullivan, Neal Trappe



Brad Gom, **David Naylor**, Locke Spencer, Ian Veenendaal



Kjetil Dohlen , Joel Lemerrer, Fabrice Madec, Eddy Rakotonimbahy, Christel Rossin, Sebastien Vives, Annie Zavagno

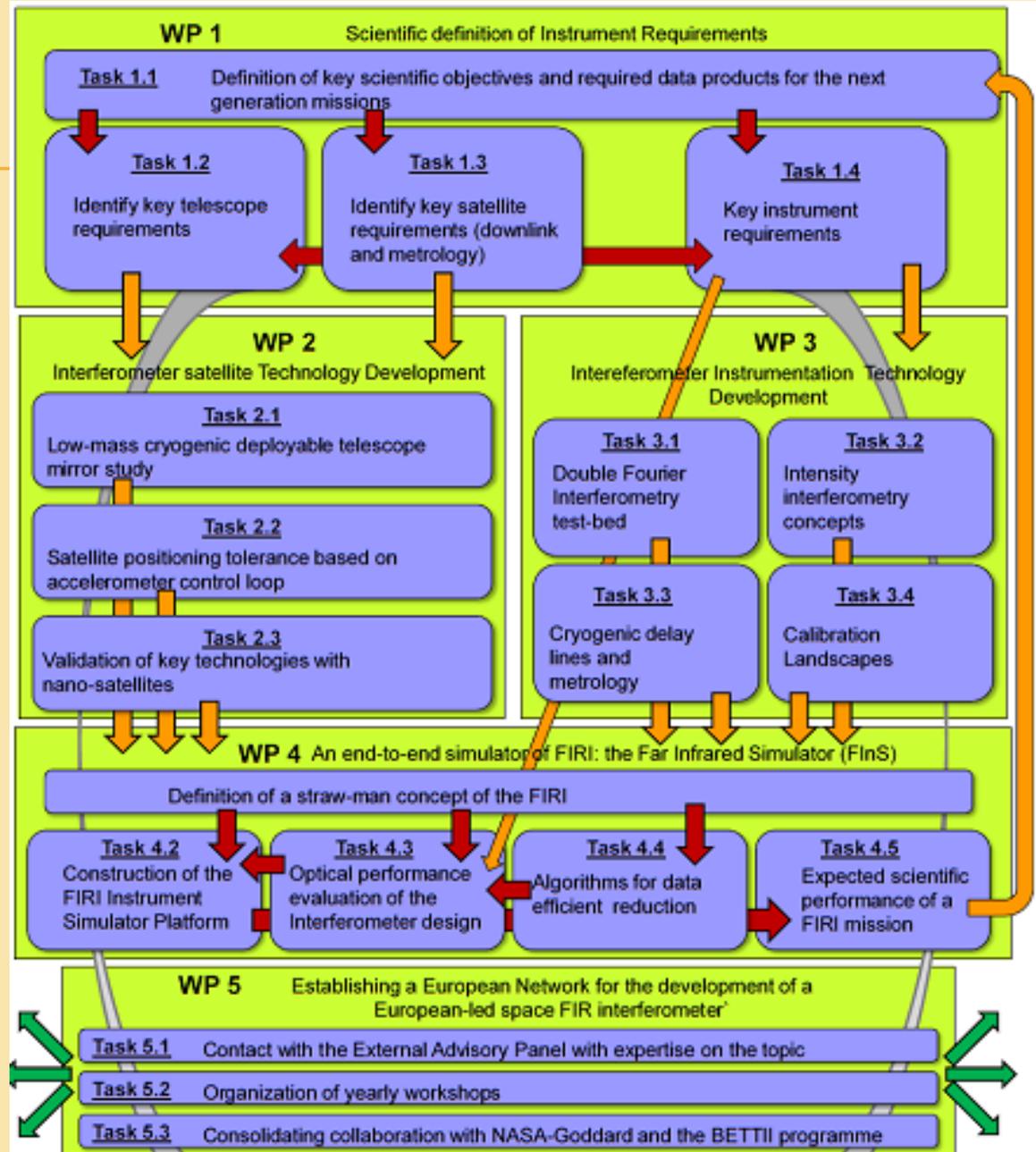


Scige' Liu, Stefano Pezzutto, **Luigi Spinoglio**.



Valerio Iafolla, Carlo Lefevre, Carmelo Magnafico, Diana Martella, Simone Pini, Daniele Schito

FISICA ACTIVITIES



FISICA ACTIVITIES

Science & Requirements
Definition

Technology Activities

Instrument Simulator
Software

Dissemination and
Networking

Science & Requirements Definition

Technology
Activities
Satellite-related

Technology
Activities
Payload-related

An end-to-end Instrument Simulator

Network and Dissemination

SCIENCE NEEDS

Near

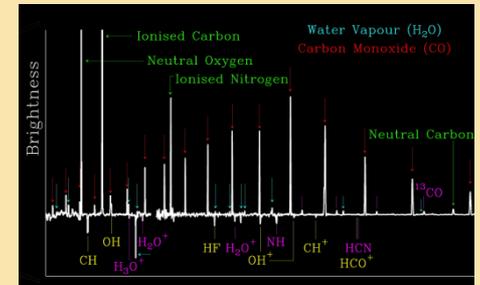
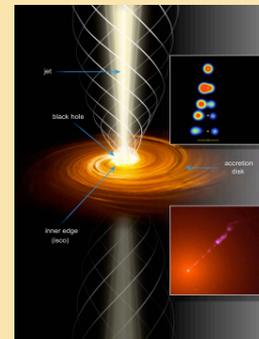
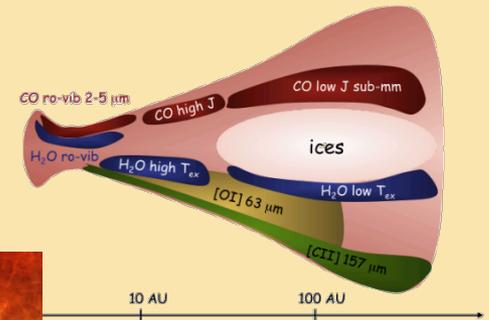
Dust and Debris-discs

Star formation

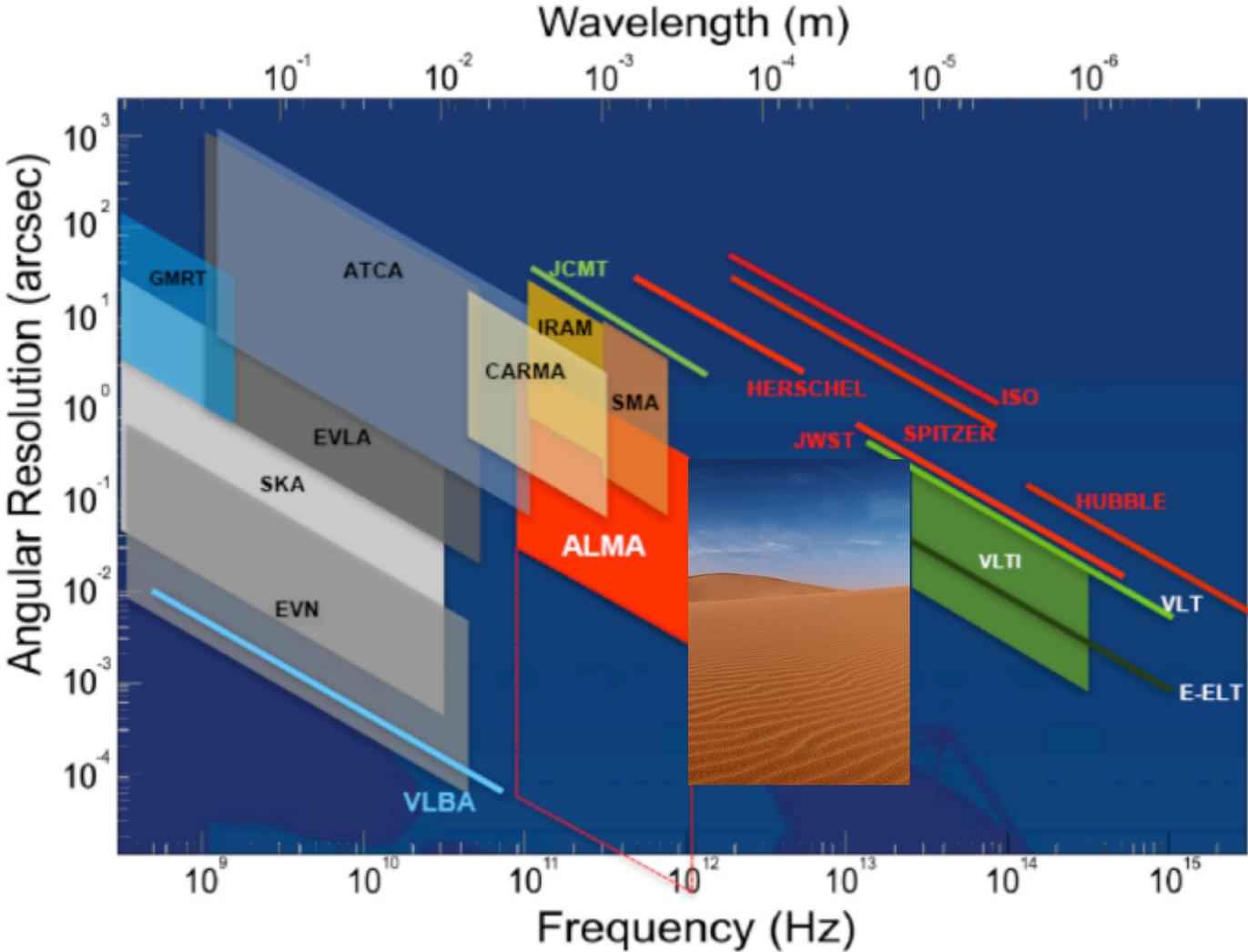
The nearby universe

The evolving universe

Far



SCIENCE IN CONTEXT



DATA PRODUCTS

Instrument Characteristics		
Primary diameter	2 m	ext
Baseline Range	0, 10-100m	ext
Optics Temperature	4K cryo-cooled	(¹)
Configuration	Rigid Truss or Tether	(³)
Mission Lifetime	3-5Yrs	ext
Sky Coverage	+/- 20 deg. from ecliptic	(^{1,2})
Spectral Coverage (μm)	25-50,50-100,100-200,200-400	(¹)
Detector Arrays (35/70/140/280 μm)	2x(28x28/14x14/8x8/4x4)	(*)
Derived parameters:		
Angular resolution	0.1 ($\lambda/100\mu\text{m}$) arcsec	(¹)
Field of View	1 arcminute	(^{1,3})
Spectral resolution	~3000-5000	(¹)ext.
Point Source Sensitivity (5s in 24hrs)	(35 / 70 / 140 / 280) μm	Ext.
-) Spectral Line (10^{-19} W/m ²)	0.7 / 0.4 / 0.3 / 0.3	
-) Continuum (μJy)	3.5 / 5 / 7.5 / 12	
Typical Observation Time	1 day	(¹)

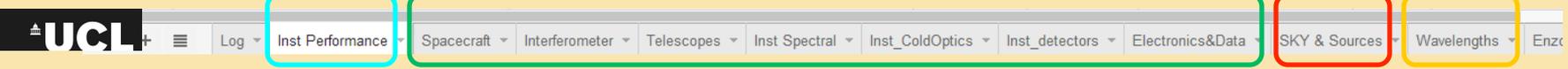
Initial numbers were based on an ideal-element sensitivity model
Consistent with a scaled version of the NASA-concept SPIRIT.

INSTRUMENT REQUIREMENTS

The Sensitivity Spreadsheet

Inter-connected sections with mutual dependencies.

Final Performances are gathered in the "performance sheet".



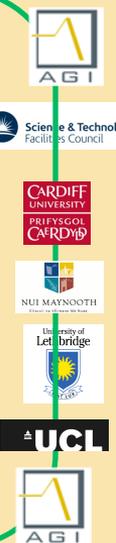
SPACECRAFT SPECIFIC PARAMETERS									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments	Dependencies	
INTERFEROMETER ARCHITECTURE PARAMETERS									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments		
DETECTORS - PARAMETERS									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments	Dependencies	
TELESCOPE PARAMETERS									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments		
INSTRUMENT SPECTRAL ARCHITECTURE DEFINITION									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments		
COLD OPTICS PARAMETERS									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments		
ELECTRONICS and DATA Parameters									
Quantity	Description	Symbol	Value	Units	Value	Units	Comments		

SKY & SOURCES PARAMETERS			
Quantity	Description	Symbol	

All numbers in columns 1 to 4 are multiplied by the overall instrument efficiency (the SSF/TensorC) has been added to the band SEDs at the end of the formulae.

Band	Blockbodies [W/cm ² /ster]				[emissivities of source included only]				SQR(Photon Flux [mW])				TOTAL - Band						
	CMB	CB	ZOZI	SUN	CMB	CB	ZOZI	SUN	CMB	CB	ZOZI	SUN	CMB	CB					
Min	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000					
Max	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000	0.0000000000000000					
λ [micrometers]	25	0.7548E			0.5 0.1 0.2 0.3 0.4	0.34E 14	2.22E 12	1.33E 10	9.90E 10	1.70E 12	4.27E 16	4.48E -01	2.48E -02	2.07E -02	1.07E -00	2.20E -02	3.09E -00	1.32E -00	5.70

INSTRUMENT PERFORMANCE QUANTITIES									
Quantity	Description	Symbol	Value	Units	Value	Units	Comment		



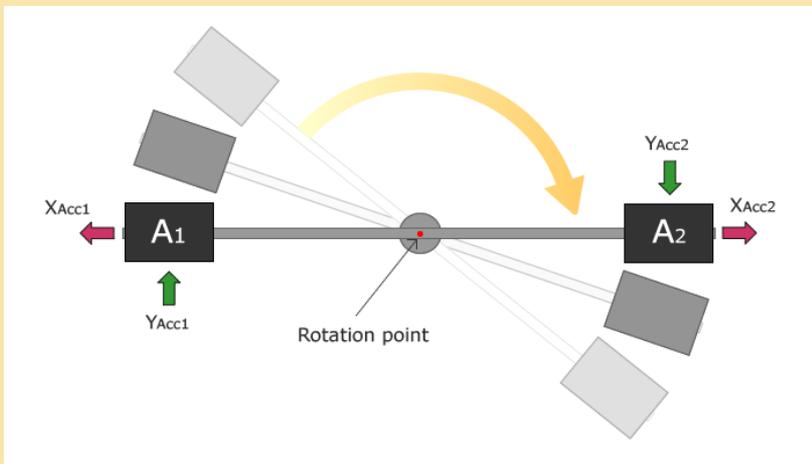
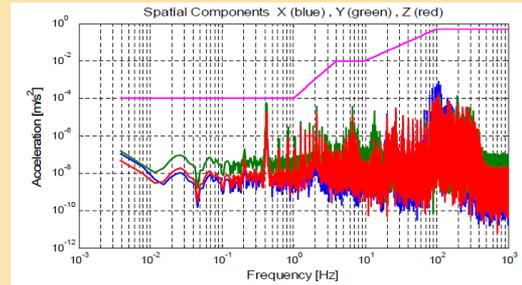
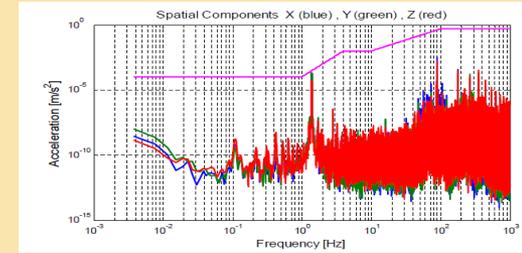
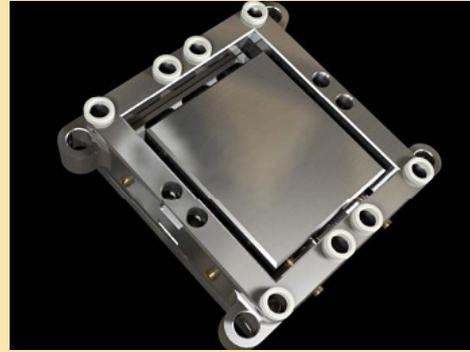
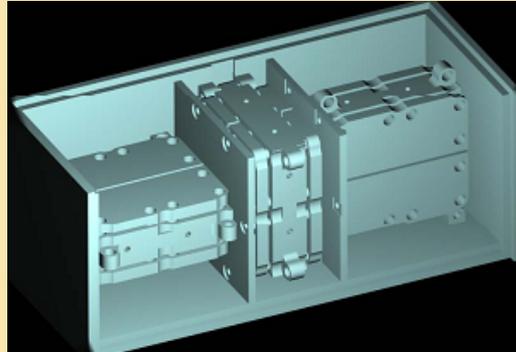
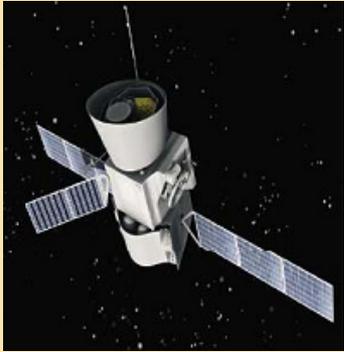
SATELLITE-RELATED ACTIVITIES

Study of tolerances and other implications of CFRPs for light-weight deployable mirrors



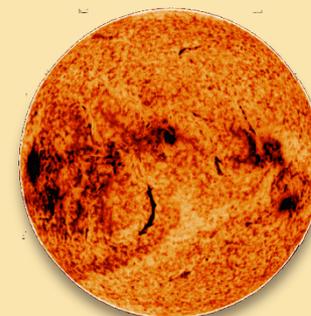
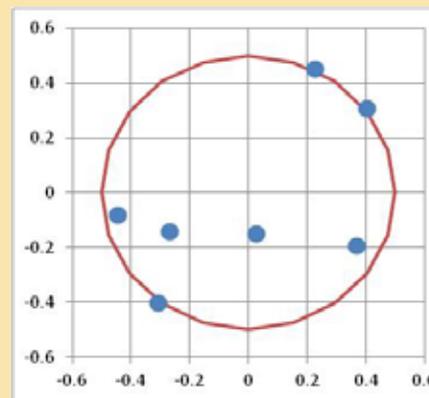
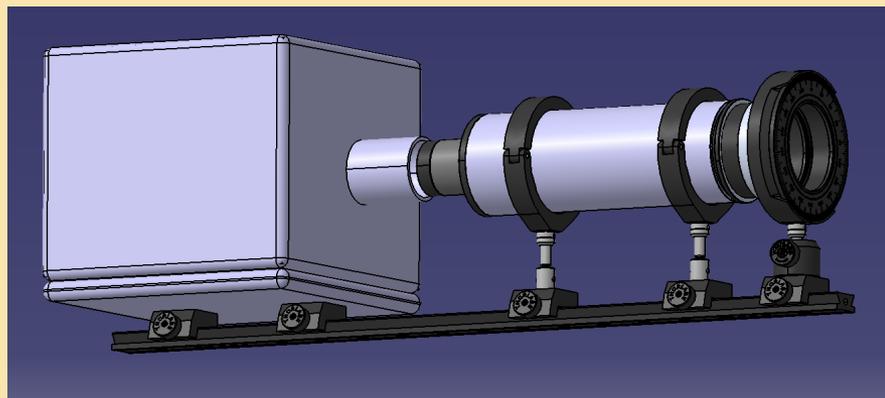
SATELLITE-RELATED ACTIVITIES

Satellite motion tolerances based on accelerometer control loop



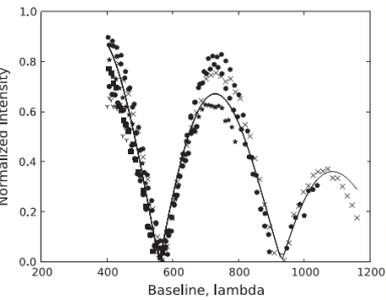
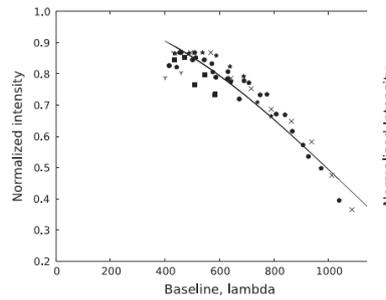
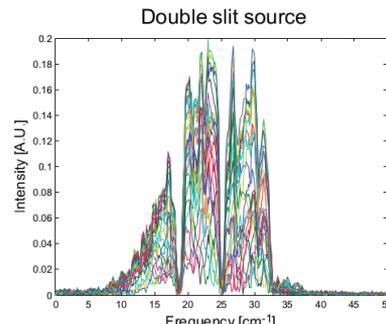
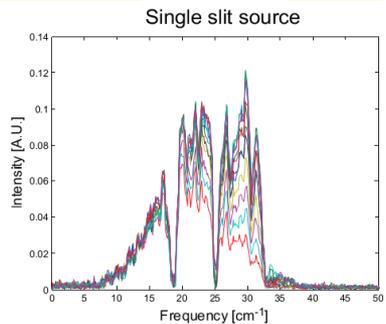
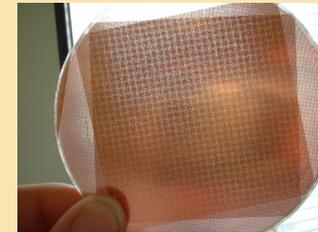
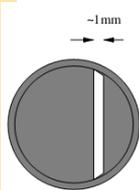
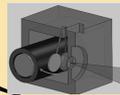
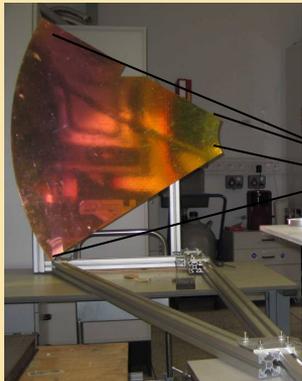
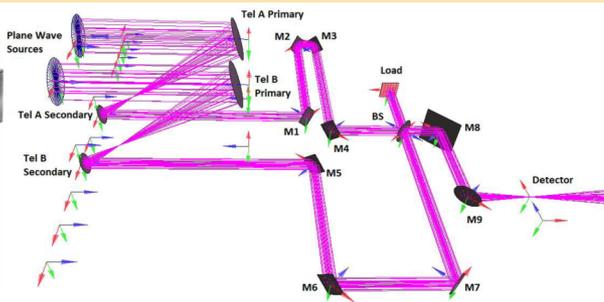
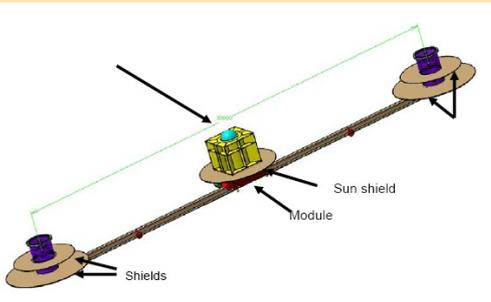
SATELLITE-RELATED ACTIVITIES

Nano-satellite test-bench validation



PAYLOAD-RELATED ACTIVITIES

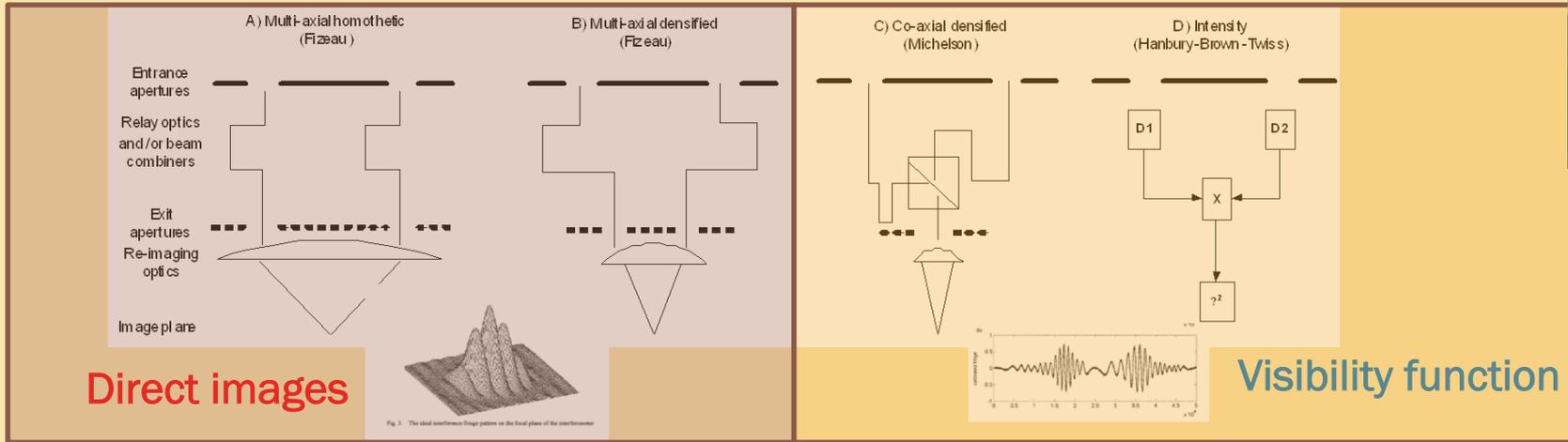
A Double-Fourier Modulation test-bed for the 30-300um range



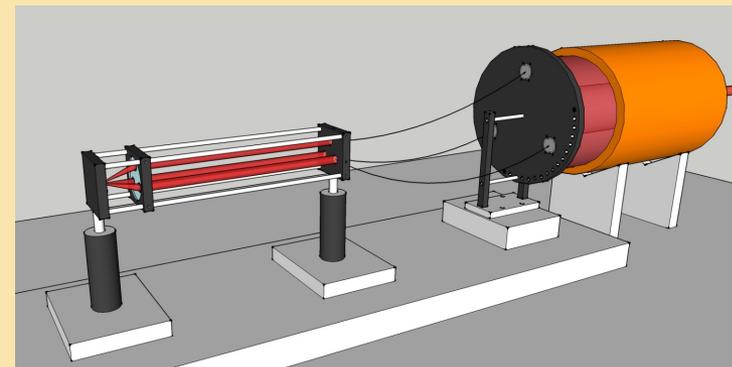
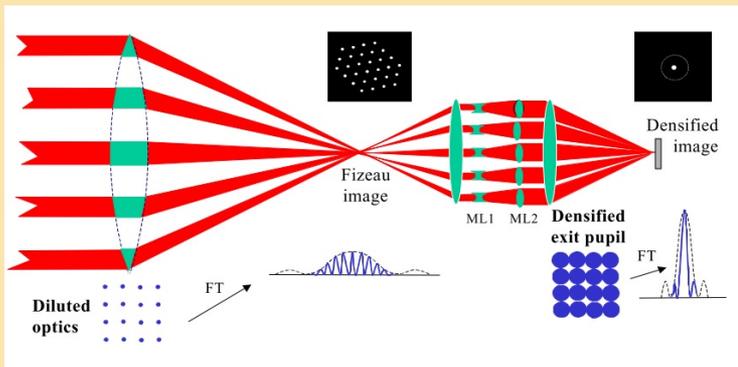
Grainger et al. 2012
FP7 Consortium

PAYLOAD-RELATED ACTIVITIES

Study of alternative techniques

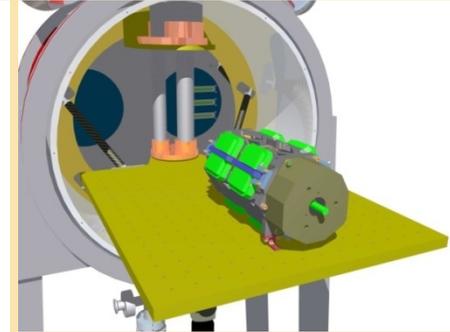
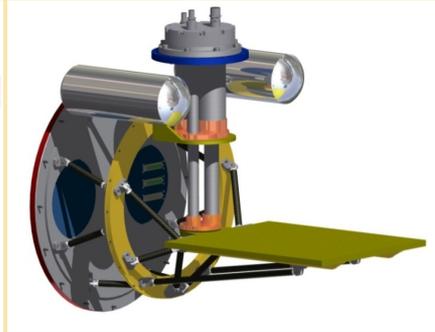
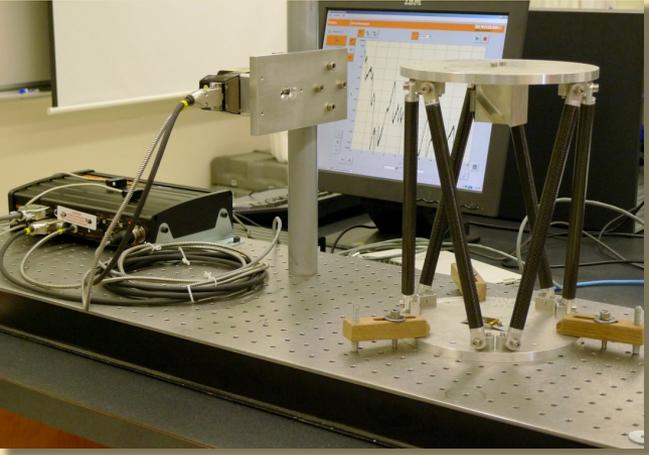


van der Avoort et al. (2007)



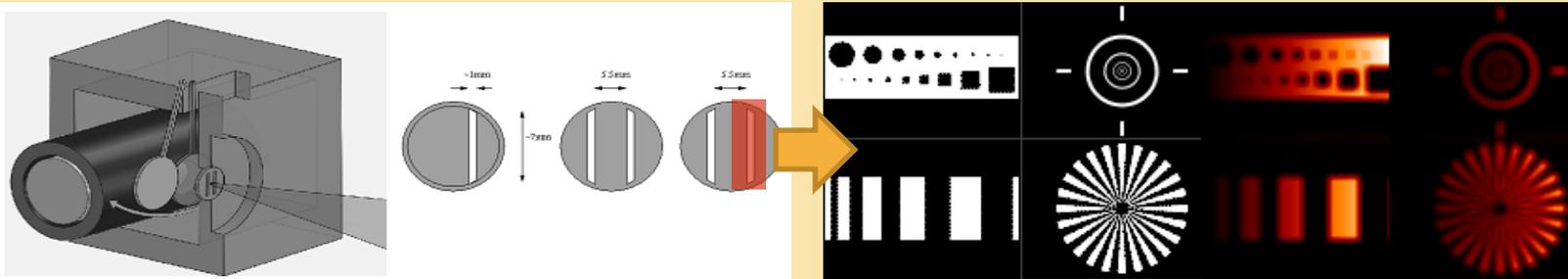
PAYLOAD-RELATED ACTIVITIES

Cryogenic (4K) delay-line metrology

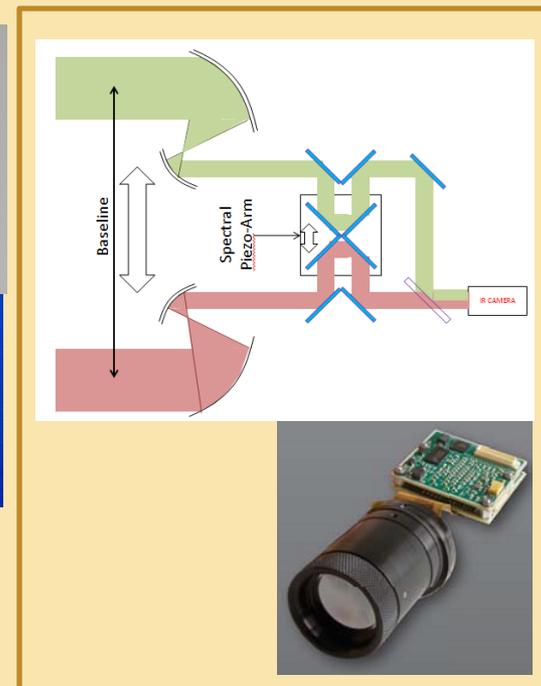
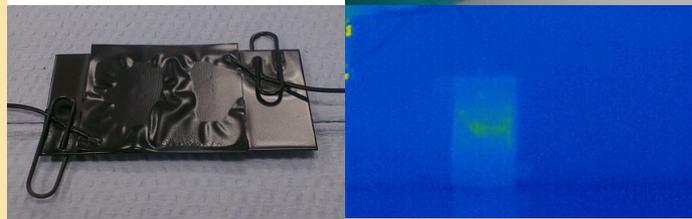
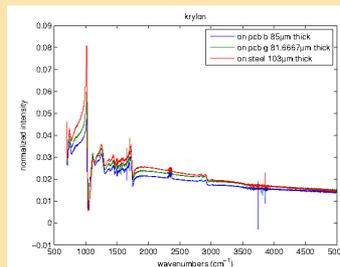
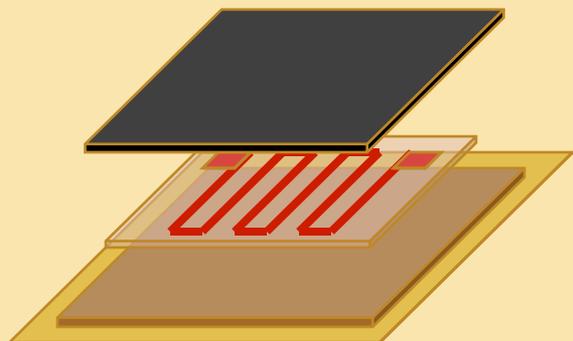


PAYLOAD-RELATED ACTIVITIES

Complex Calibration Sources

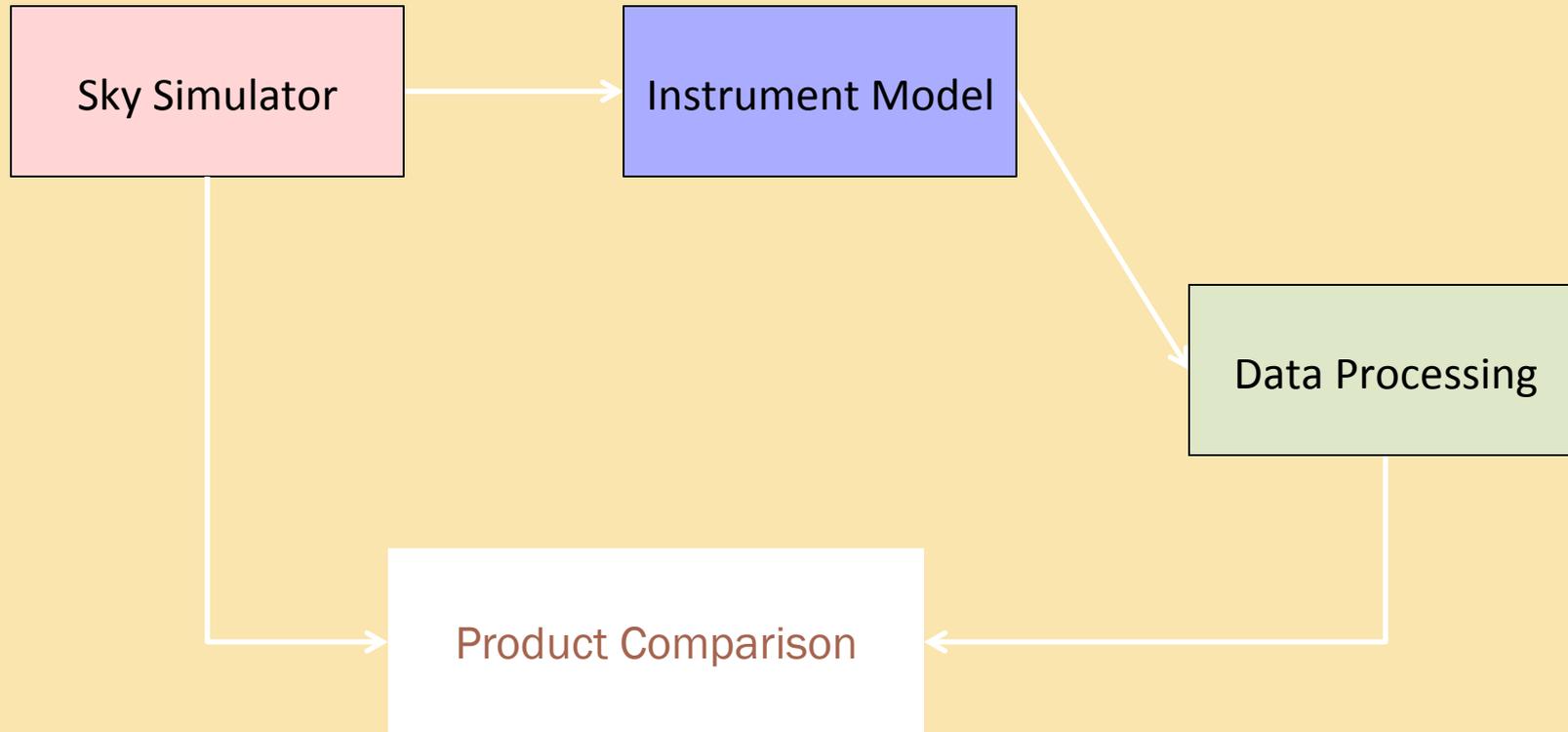


Polypropilene based copper dissipator



FIINS (FAR INFRARED INSTRUMENT SIMULATOR) - 1

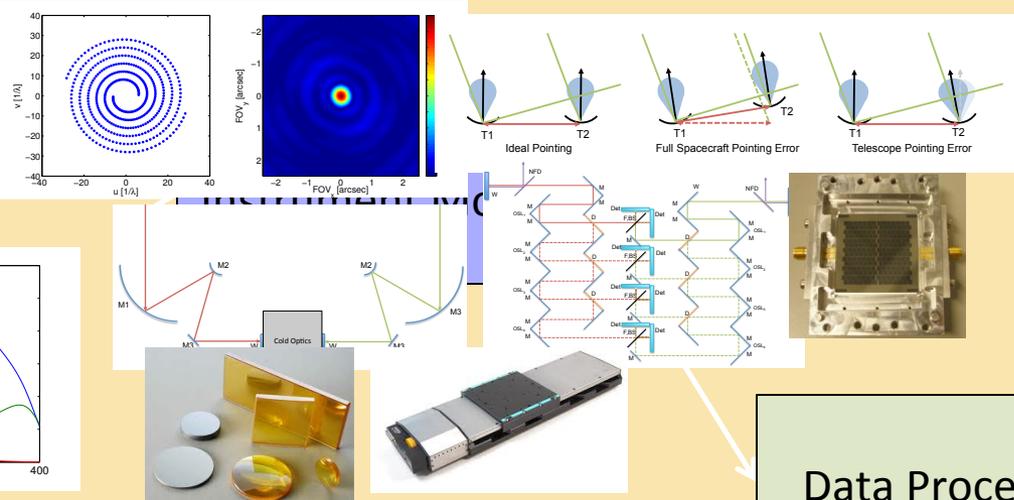
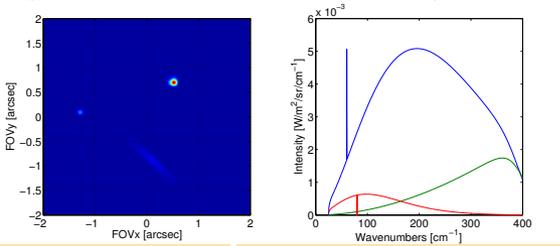
Based on the original work of Roser Juanola-Parramon (Doctoral Thesis)



FIINS (FAR INFRARED INSTRUMENT SIMULATOR) - 1

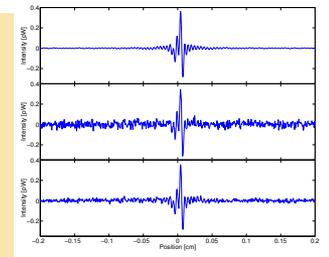
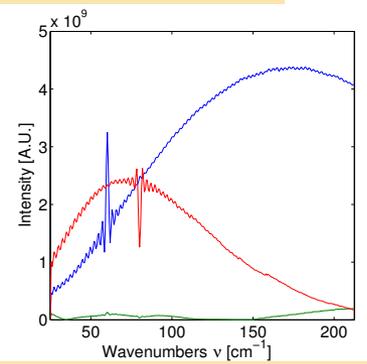
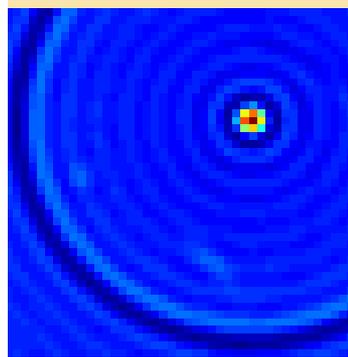
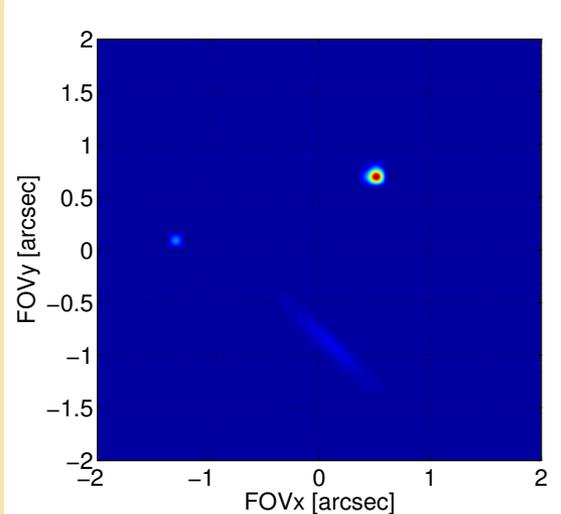
Based on the original work of Roser Juanola-Parramon (Doctoral Thesis)

Sky Simulator



Data Processing

$$I_g(\delta, b) = I_1 + I_2 + 2 \sum_{i=0}^{N_{FOV}} \sum_{j=0}^{N_{FOV}} \sum_{k=0}^N |SkyC(\theta_{x,i}, \theta_{y,j}; \nu_k)| \cos[2\pi\nu_k\delta - 2\pi\nu_k(\mathbf{b} \cdot \boldsymbol{\theta})]$$



EU - NETWORK

DISSEMINATION

WP 5 European Network creation for the development of a European-led space FIR interferometer

Task 5.1 Contact with the External Advisory Panel with expertise on the topic

Task 5.2 Consolidating collaboration with NASA-Goddard and the BETTII programme

WP 6 Dissemination of scientific results

Task 6.1 Organization of yearly workshops

Task 6.2 Participation to conferences



EU - NETWORK

DISSEMINATION

Yearly workshop as a reference and to allow focused approach

Bringing Fundamental Astrophysical Processes Into Focus: A Community Workshop to Plan the Future of Far-Infrared Space Astrophysics, Goddard

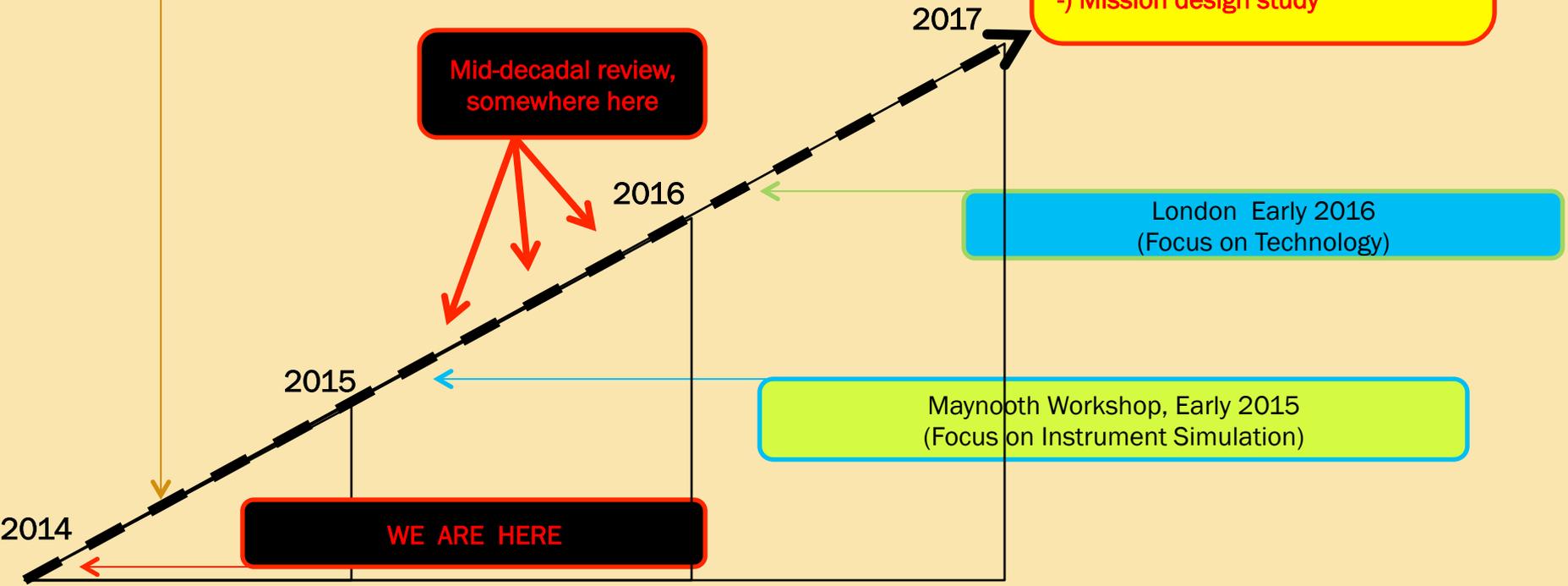
The next opportunities:
-) pathfinders
-) technology demonstrators
-) Mission design study

Mid-decadal review, somewhere here

London Early 2016
(Focus on Technology)

Maynooth Workshop, Early 2015
(Focus on Instrument Simulation)

WE ARE HERE



CONCLUSIONS

- 2013 has been an exciting (not all good) year for the far-infrared
 - “Death” of Herschel, but not of its science
 - L2/L3 call – pooling of ideas and resources
 - The SPICA next step...

- FISICA-FP7 will focus on the identification and definition of the key data products required from the science + analysis of the requirements of a space interferometer to achieve these.

- Technological activities relevant to satellite and instrument have commenced

- While the program cannot be comprehensive of all techniques and existing technology groups interested in the FIR, the Networking and Dissemination elements allow for a regular note-comparing exercise in order to keep focus

- For information on the Consortium (and future related workshops):
www.fp7-fisica.eu

- There is a plan (not yet implemented) to use the webpage as a repository of Far-Infrared science and mission concepts to allow the community to access relevant information as required.



**THANKS FOR
YOUR KIND
ATTENTION**

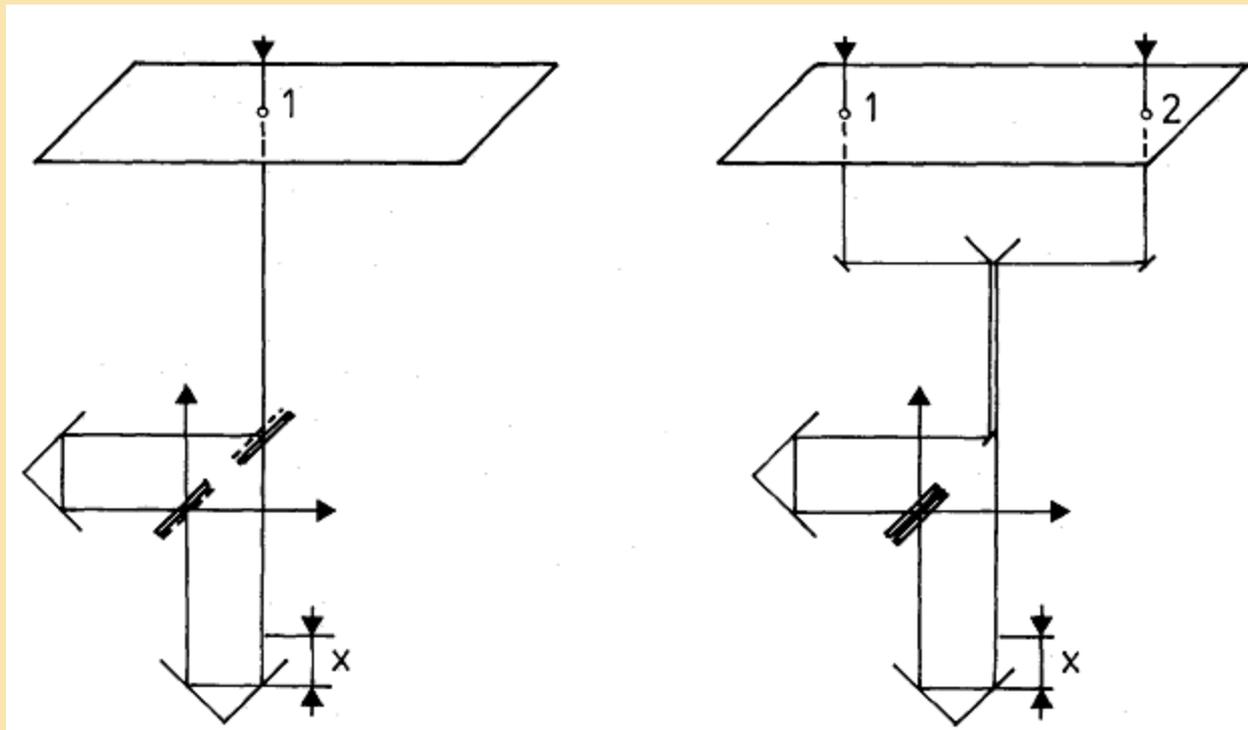
ADDITIONAL MATERIAL

DOUBLE-FOURIER MODULATION (DFM)

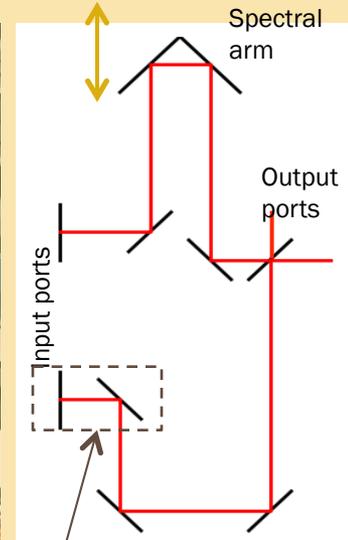
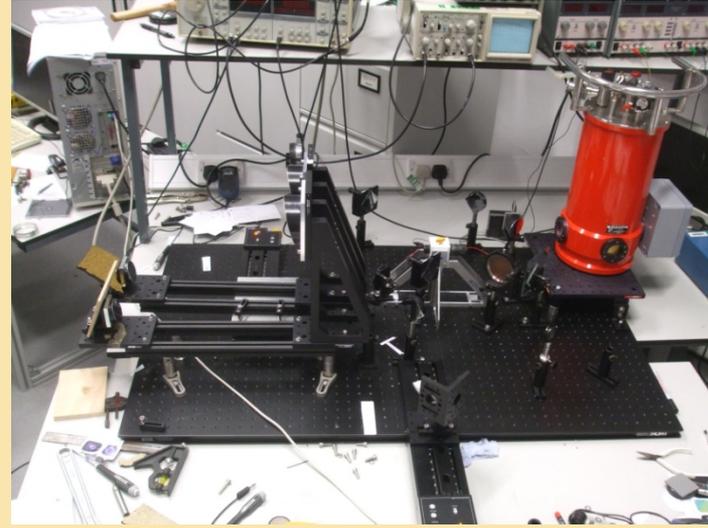
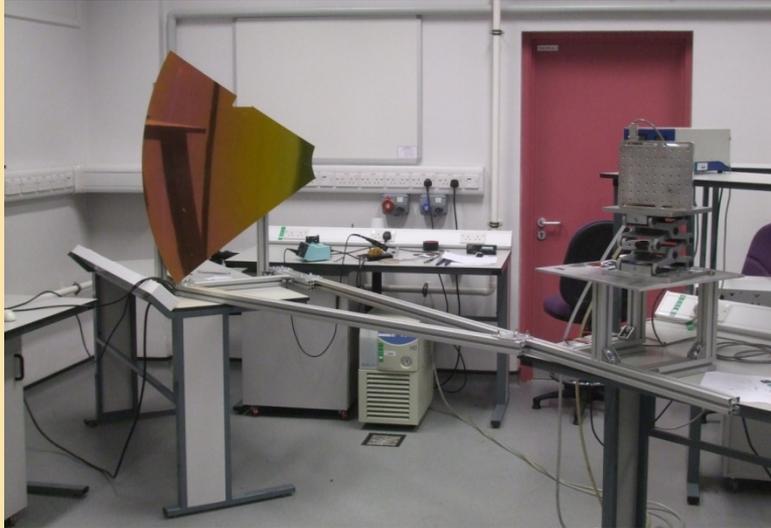
Astron. Astrophys. 195, 350–363 (1988)

Double Fourier spatio-spectral interferometry: combining high spectral and high spatial resolution in the near infrared *

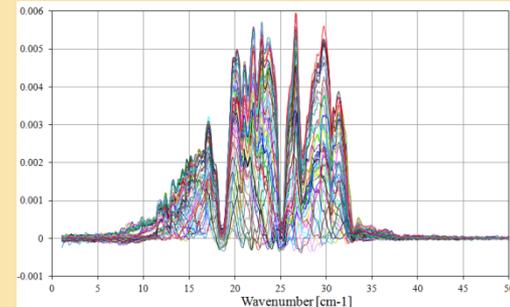
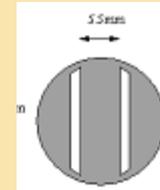
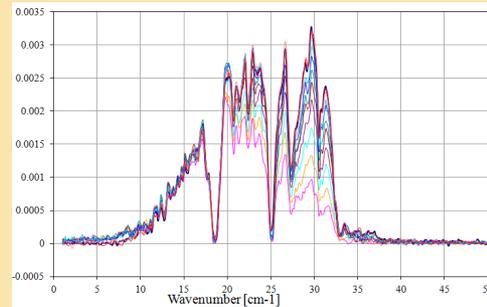
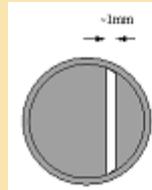
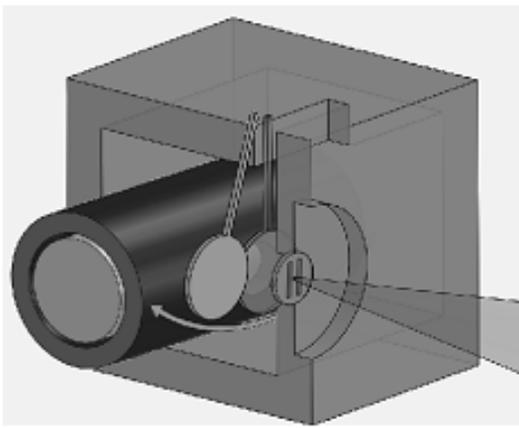
J.-M. Mariotti¹ and S.T. Ridgway²



A TEST-BED FOR THE FIRI OPTICS DEVELOPMENT

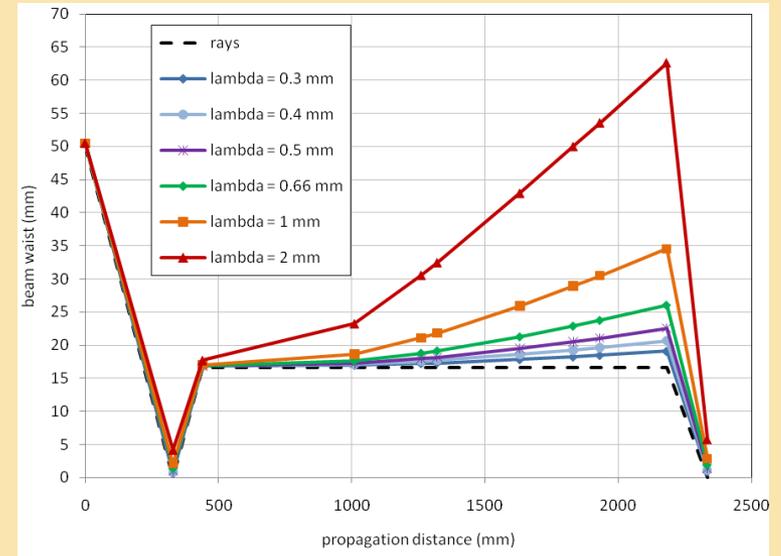
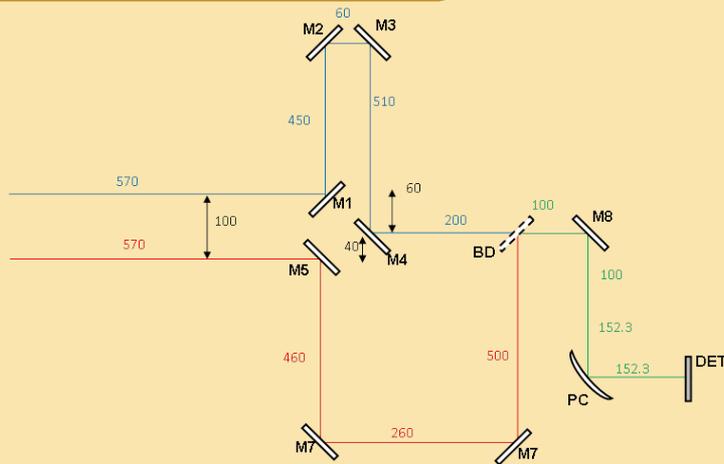


Variable baseline

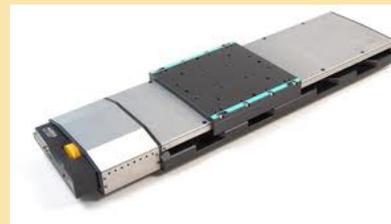


TEST-BED UPGRADES PLANNED - MID & FAR INFRARED

Detailed optical modelling

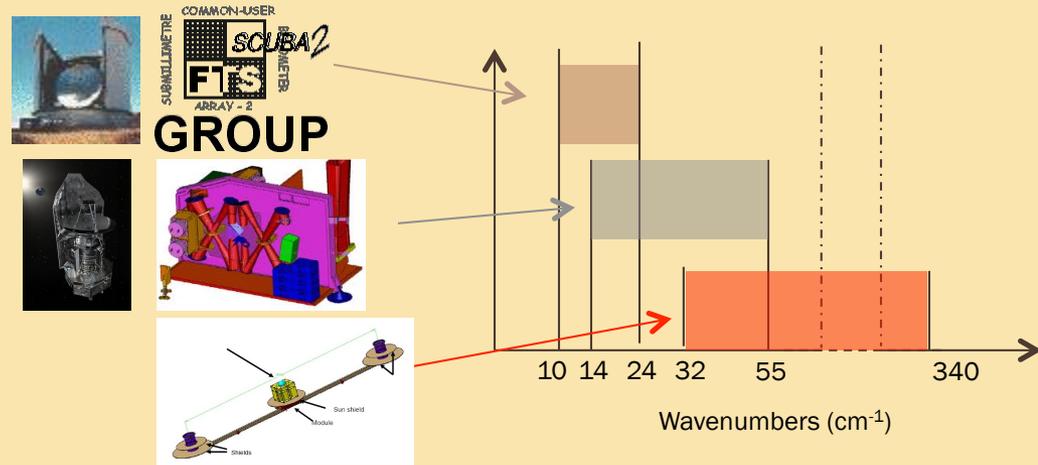


Upgrade metrology

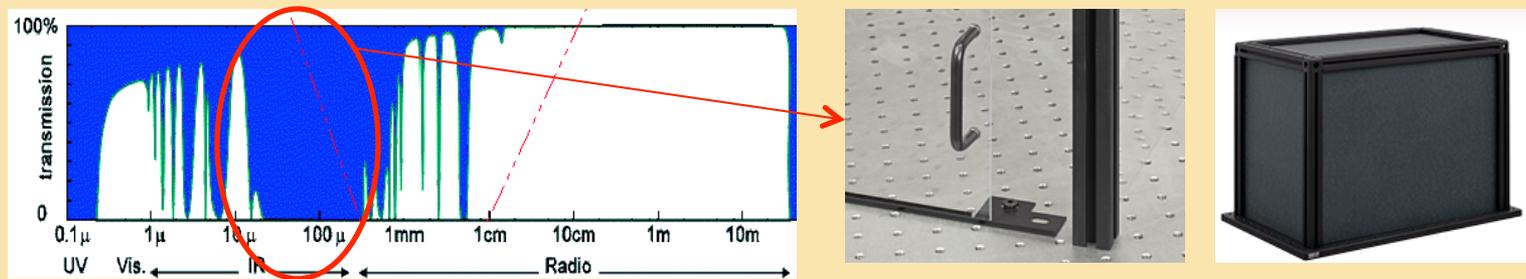


TEST-BED UPGRADES PLANNED - MID & FAR INFRARED

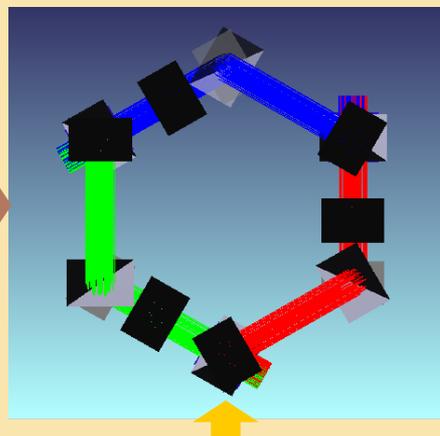
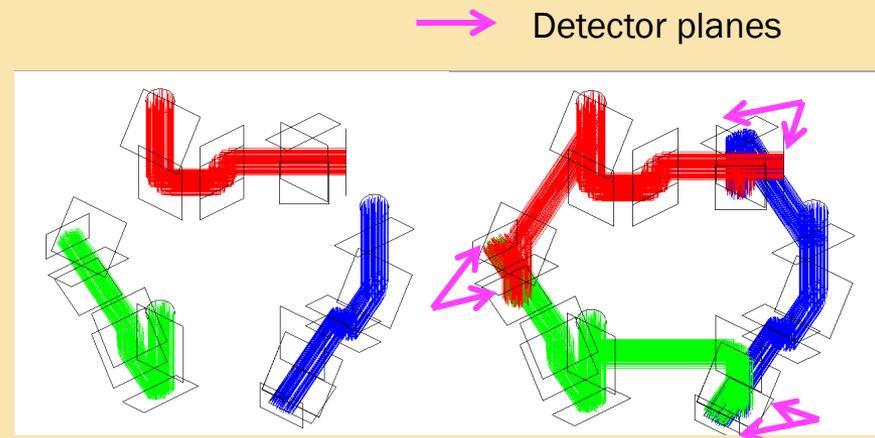
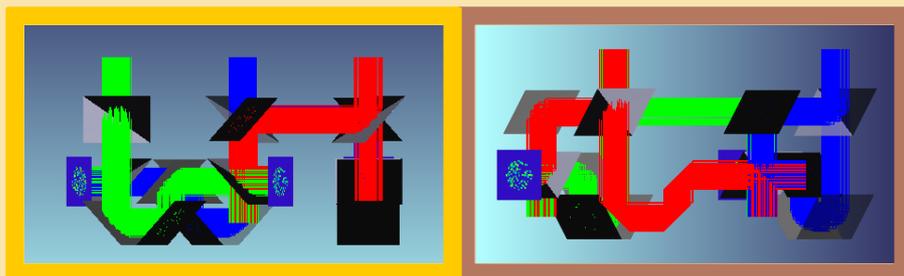
A wide-band beam splitter
(or maybe two)



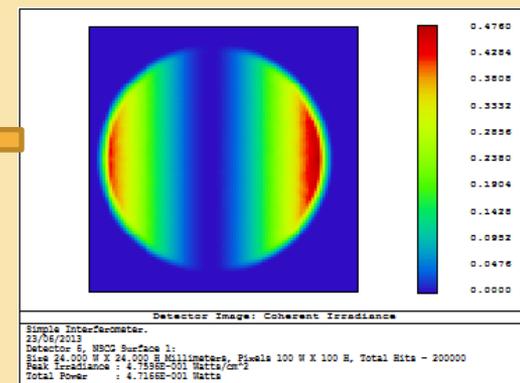
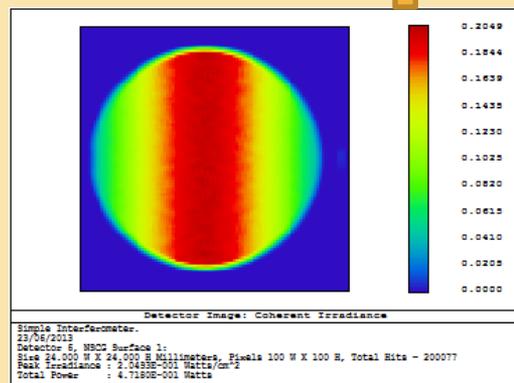
Optical bench conversion



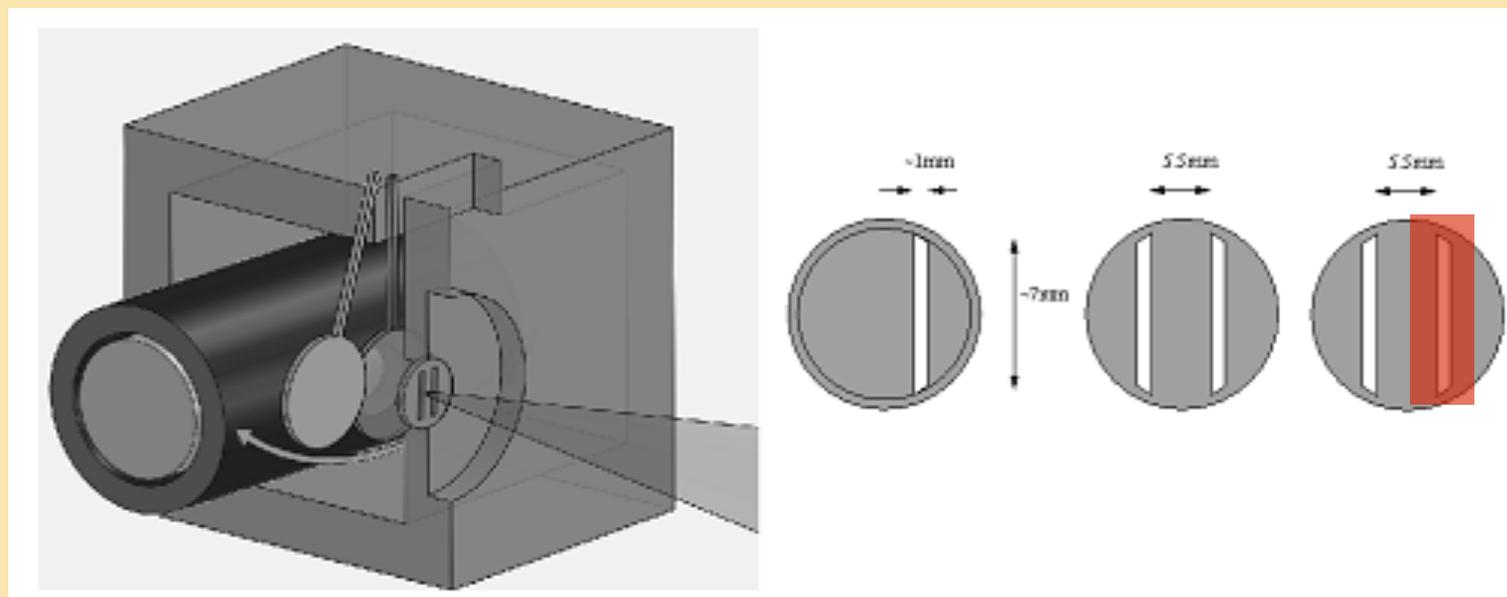
TEST-BED UPGRADES PLANNED – 3 BEAMS?



With additional
K-mirrors for
image rotation.

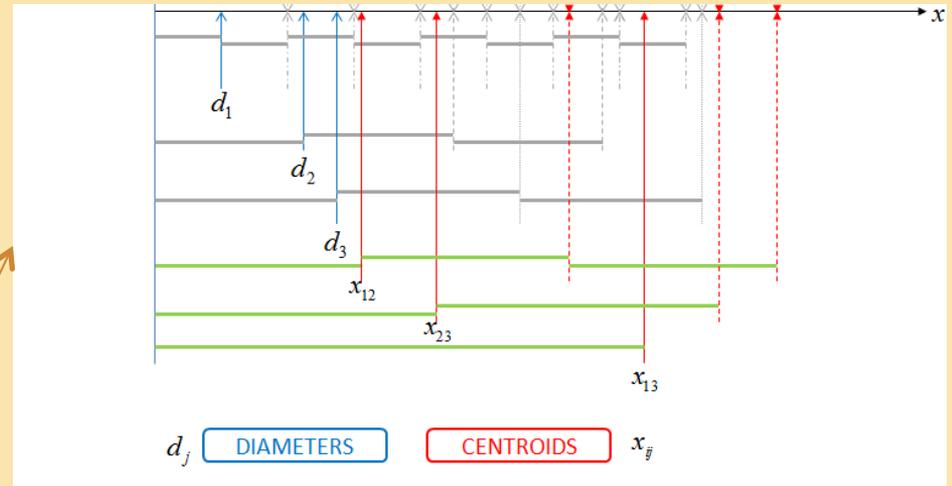
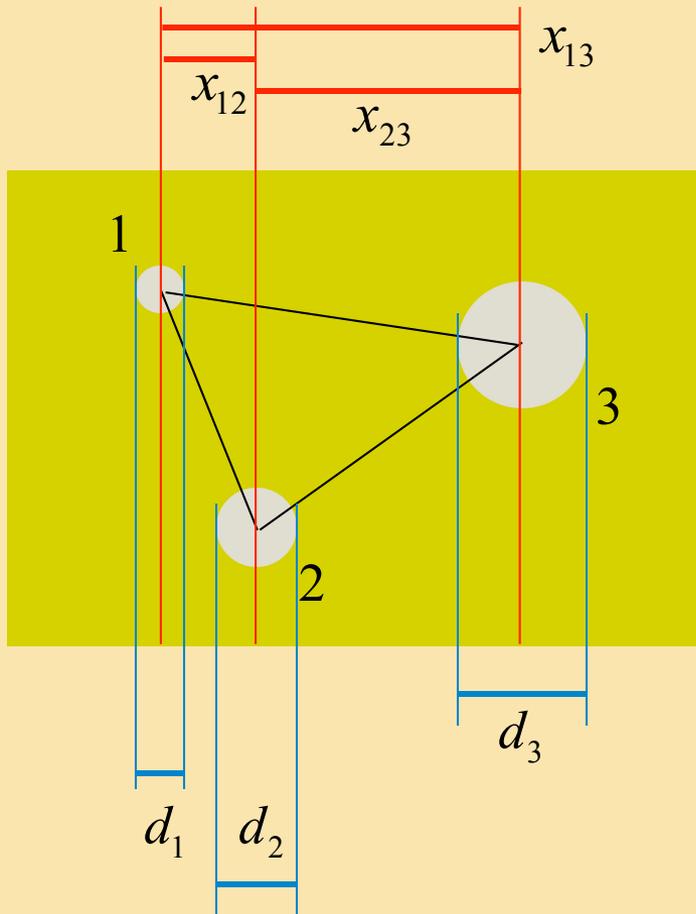


CALIBRATION SOURCES (FROM SIMPLE TO COMPLEX)

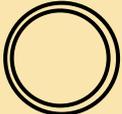
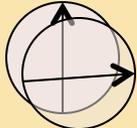


Grainger et al. 2012

CALIBRATION SOURCES (FROM SIMPLE TO COMPLEX)



Potentially covered by

-  Filters
-  Diaphragms
-  Variable FP-etalon

CALIBRATION SOURCES (FROM SIMPLE TO COMPLEX)

